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Journal of the Atmospheric Sciences

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The DOI for this manuscript is doi: 10.1175/JAS-D-13-0190.1

The final published version of this manuscript will replace the preliminary version at the above DOI once it is available.

If you would like to cite this EOR in a separate work, please use the following full citation:

Zhou, W., I. Held, and S. Garner, 2013: Parameter study of tropical cyclones in rotating radiative-convective equilibrium with column physics and resolution of a 25 km GCM. J. Atmos. Sci. doi:10.1175/JAS-D-13-0190.1, in press.

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## Parameter study of tropical cyclones in rotating 1 radiative-convective equilibrium with column physics and 2 resolution of a 25 km GCM 3 WENYU ZHOU 4 Atmospheric and Oceanic Sciences Program, Princeton University, Princeton, New Jersey ISAAC M. HELD 5 Geophysical Fluid Dynamics Laboratory/NOAA, Princeton, New Jersey STEPHEN T. GARNER Geophysical Fluid Dynamics Laboratory/NOAA, Princeton, New Jersey RELIMIN

#### ABSTRACT

<sup>7</sup> Rotating radiative-convective equilibrium is studied by extracting the column physics of a <sup>8</sup> meso-scale resolution global atmospheric model that simulates realistic hurricane frequency <sup>9</sup> statistics and coupling it to rotating hydrostatic dynamics in doubly-periodic domains. The <sup>10</sup> parameter study helps in understanding the tropical cyclones simulated in the global model <sup>11</sup> and also provides a reference point for analogous studies with cloud resolving models.

The authors first examine the sensitivity of the equilibrium achieved in a large square domain  $(2 \times 10^4 \text{ km} \text{ on a side})$  to sea surface temperature, ambient rotation rate and surface drag coefficient. In such a large domain, multiple tropical cyclones exist simultaneously. The size and intensity of these tropical cyclones are investigated.

The variation of rotating radiative-convective equilibrium with domain size is also studied. As domain size increases, the equilibrium evolves through four regimes: a single tropical depression, an intermittent tropical cyclone with intensity widely varying, a single sustained storm, and finally multiple storms. As SST increases or ambient rotation rate f decreases, the sustained storm regime shifts towards larger domain size. The storm's natural extent in large domains can be understood from this regime behavior.

The radius of maximum surface wind, although only marginally resolved, increases with SST and increases with f for small f when the domain is large enough. But these parameter dependencies can be modified or even reversed if the domain is smaller than the storm's natural extent.

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### <sup>26</sup> 1. Introduction

Rotating radiative-convective equilibrium, achieved in a doubly periodic box on an f-plane with horizontally homogeneous forcing and boundary condition, is an informative idealized framework for studying the interactions between moist thermodynamics, radiation, and rotating dynamics. It can be studied in non-hydrostatic models in which deep convection is partly resolved and also in lower resolution hydrostatic models with parameterized convection. In the latter case, one can think of this framework as one of a hierarchy of idealized settings in which to study the implications of the assumptions made in GCMs.

Tropical cyclones (TCs) form within these rotating radiative-convective equilibria, and 34 they provide simulations in which very long-lived mature TCs emerge. If run at the resolu-35 tion and with the column physics (including boundary layer, radiation, microphysics, moist 36 convection and cloud modules) of a global comprehensive model, this idealized framework 37 can help us evaluate the TC simulations in such models. Global models are moving to high 38 enough resolution that aspects of their TC simulations are becoming more realistic, and 39 these global models are one of the tools being used to try to predict the impact of climate 40 change on TC statistics. Being able to study the model-generated TCs in this idealized ge-41 ometry, eventually comparing to analogous simulations at cloud-resolving resolutions, should 42 provide valuable information as to the limitations and strengths of the TC simulations in 43 global models. 44

Held and Zhao (2008, hereafter HZ08) describe such a rotating radiative-convective equi-45 librium in which the column physics of a particular atmospheric model, GFDL's AM2 (At-46 mospheric Model version 2) is coupled to rotating hydrostatic dynamics in a doubly-periodic 47 box. A similar equilibrium is achieved in Nolan et al. (2007) and Nolan and Rappin (2008) 48 with higher cloud-resolving resolution in a study of tropical cyclogenesis. Schecter and 49 Dunkerton (2009) and Schecter (2011) set up an idealized three-layer hurricane model under 50 such a framework, with simplified radiative and convective parameterizations. The variabil-51 ity and predictability of hurricanes in rotating radiative-convective equilibrium is studied by 52

Hakim (2011), Hakim (2013), and Brown and Hakim (2013) using both axisymmetric and
three-dimensional models.

Besides their various choices for column physics and resolution, a major difference among 55 these studies is that HZ08 considers a sufficiently large domain for multiple TCs to coexist 56 while others use small domains within which only one TC develops. In those small-domain 57 simulations, whether with a doubly-periodic domain or a rigid-wall box, the extent of TCs 58 is explicitly set by the domain size. Because of this constraint, both the state of equilibrium 59 achieved and its parameter sensitivity may be interfered with by the domain size. We refer 60 to our doubly-periodic small-domain simulation as *lattice* equilibrium, in which duplicated 61 TCs exist simultaneously at the grid points in a lattice that can be visualized by duplicating 62 the doubly-periodic domain along its boundaries. The large-domain simulations avoid such 63 finite-domain effects and can be regarded as *turbulent* equilibria in the sense that multiple 64 TCs interact with respect to each other in more complex ways. The extent of TCs is thus 65 controlled by internal processes of the turbulent equilibrium instead of domain size. This 66 allows for analysis of size and intensity of TCs as a function of environmental parameters, 67 independent of domain size. Neither configuration is realistic in the sense that appropriate 68 decay mechanisms (poleward drift and interactions with land and extratropical systems) 69 and suppression factors (wind shear, subsidence and dry air entrainment) are not present. 70 resulting in very long-lived storms. 71

In this study, both lattice and turbulent equilibria are achieved in doubly-periodic do-72 mains, with column physics and resolution of a 25 km global atmospheric model, which 73 simulates many aspects of tropical cyclone statistics realistically (Zhao et al. 2009). The 74 details of model formulation are presented in section 2. We start by describing turbulent 75 equilibrium achieved in a large domain  $(2 \times 10^4 \text{ km on a side})$  in section 3. The parameter 76 sensitivity of TC structure and intensity in this turbulent equilibrium is studied by varying 77 sea surface temperature, ambient rotation rate f and surface drag coefficient  $C_D$ . In section 78 4, we explore the variation of lattice equilibrium as a function of domain size, and show how 79

the natural extent of TCs in turbulent equilibrium is related to this domain size dependency
of lattice equilibrium. A discussion and conclusion are offered in section 5.

One of our goals is to map out some of these parameter dependencies in the hope of en-82 couraging the study of other GCM's column physics in this idealized setting. We believe that 83 understanding these parameter dependencies and how they differ between parameterized-84 convection hydrostatic and cloud-resolving non-hydrostatic models will be helpful in eval-85 uating the simulations of TC statistics in global models in the future. To the extent that 86 reliable cloud-resolving simulations are available, this idealized rotating radiative-convective 87 equilibrium framework could also be used as a novel test for the performance of convection 88 schemes. 89

### <sup>90</sup> 2. Model formulation

The column physics used here is that of HiRAM (GFDL High Resolution Atmospheric 91 Model). HiRAM has been used in Zhao et al. (2009) to realistically simulate the observed 92 global climatology and interannual variability of hurricane frequency at 50 km resolution, 93 although few storms are simulated higher than category 2 as measured by the maximum 94 surface winds, especially in the N.Atlantic (Zhao and Held 2010). Multiple simulations of 95 the period 1979-2008, running over observed sea surface temperatures, with both 50 km and 96 25 km resolution versions of this model have been placed in the CMIP5 (Climate Model 97 Intercomparison Project Version 5) archive. The seasonal and interannual variability of 98 tropical storm genesis is not fundamentally different in the 25 and 50 km models. A 25 km 99 version of the model with modified surface flux formulation and microphysical scheme has 100 shown impressive skill for seasonal forecasts in the Atlantic (Chen and Lin 2012). 101

HiRAM was developed starting from an earlier lower resolution model, AM2 (GAMDT
 2004). The prognostic cloud fraction scheme is changed to a simpler diagnostic scheme
 assuming a sub-grid scale distribution of total water. The relaxed Arakawa-Schubert con-

vective closure (Moorthi and Suarez 1992) in AM2 has been replaced by a modified shallow 105 convection scheme (Bretherton et al. 2004 and Zhao et al. 2009), which results in a larger 106 fraction of precipitation occurring at resolved scale rather than through the parameteriza-107 tion. All other physical schemes are the same as in AM2. Surface fluxes are computed using 108 Monin-Obukhov similarity theory, with a gustiness component  $(1 \text{ m s}^{-1})$  to account for 109 contribution from sub-grid wind fluctuation (Beljaars 1995). Oceanic roughness lengths are 110 prescribed according to Beljaars (1995). For vertical diffusion, a K-profile scheme designed 111 for cloud-topped boundary layers, with parameterized entrainment rates at the top of the 112 boundary layer (Lock et al. 2000) is used. Stability functions with thresholds dependent on 113 Richardson number are adopted for stable layers. We refer the reader to GAMDT (2004) 114 and Zhao et al. (2009) for further details. 115

The doubly-periodic model has the same vertical level altitudes as the GCM with the lowest level at about 35 m. The horizontal resolution is 25 km. As in HZ08, there is no diurnal or seasonal cycle in the doubly periodic model, and the radiative forcing is configured using an equatorial annual mean solar zenith angle. Sea surface temperatures are homogeneously prescribed over a saturated (ocean) surface. Aerosols are absent in the atmosphere, while the stratospheric ozone is fixed at an observed tropically-averaged vertical profile.

The homogeneous doubly-periodic f-plane framework contains none of the synoptic vari-122 ability that helps control TC genesis in the HiRAM global models, but the genesis frequency 123 in those models is not simply a product of downscale transfer from larger scales; it is also 124 a function of parameters in the convection scheme (Zhao et al. 2012). As one example, if 125 one suppresses deep convection very strongly by increasing the strength of entrainment into 126 the parameterizations deep convective cores, the number of TCs that form in the global 127 model is reduced dramatically. Very similar behavior is found in our homogeneous f-plane 128 simulations. We hope to describe the dependence of rotating radiative-convective equilibria 129 at this resolution on the convection scheme elsewhere. 130

# <sup>131</sup> 3. Varying environmental parameters in turbulent equi <sup>132</sup> librium

Turbulent equilibrium is achieved using a large domain  $(2 \times 10^4 \text{ km})^2$ . The initial con-133 dition is an observed tropically-averaged temperature profile, with moisture artificially near 134 saturation below 100 hPa. The model is integrated 2.5 years to its equilibrium. Most statis-135 tics, such as hydrological balance, do not require this long an integration. After the initial 136 genesis period a qualitative equilibrium is typically achieved in a few months, but the number 137 of TCs usually decreases slowly with time until it is finally stable. While we cannot prove 138 that the TC number would not decrease with even longer integrations, the behavior seen in 139 the last 6 months of the simulations is a statistically stationary number of cyclones, with 140 the occasional collapse or merger event followed by genesis into the space left by this event. 141 For every 3-hourly output during the last 6 months of each simulation, TCs are collected by 142 identifying points with surface pressure  $p_s$  satisfying two criteria: 1)  $p_s$  is less than a critical 143 value of 990 hPa, and 2)  $p_s$  is a minimum at this point within the surrounding 10×10 grid 144 box. 145

We first look at the sensitivity to SST, with ambient rotation rate set at 20°N. Snapshots 146 of surface wind speed for two different SSTs are compared in Fig. 1. As found in HZ08, the 147 number of TCs decreases with SST. There are about 40 TCs within the domain for SST at 297 148 K but only about 15 TCs for SST at 305 K. The distribution of TC central surface pressure is 149 shown in Fig. 2a. The peak frequency of TC central surface pressure shifts towards a lower 150 value as SST increases, indicating consistent higher intensity at warmer SST. The mean 151 central surface pressure intensifies by about 3 hPa  $K^{-1}$  and the mean maximum surface 152 wind speed (the wind in the lowest model layer, at roughly 35 m) increases roughly 1.5 m 153  $s^{-1}$  K<sup>-1</sup> (Fig. 3a and 3b). This is comparable to the sensitivity found in the global model, 154 which indicates roughly a  $4 \text{ m s}^{-1}$  increase in the mean maximum surface wind between the 155 case with climatological SST and that with 2 K increases in SST (Zhao and Held 2010). 156

Potential intensity (PI) is computed from the time- and domain-mean thermodynamic 157 profiles, following Bister and Emanuel  $(2002a)^{1}$ . We choose the option of pseudo-adiabatic 158 ascent and include dissipative heating (as in the GCM, the kinetic energy dissipation in the 159 boundary layer is returned as heat within the boundary layer). The ratio of surface exchange 160 coefficient for enthalpy  $C_k$  to surface drag coefficient  $C_D$  is set at 0.6, to be consistent with 161 the ratios produced in the model. A velocity reduction factor of 0.9 is applied as a way to 162 extrapolate down through the frictional boundary layer (Bister and Emanuel 2002b). The 163 theoretical maximum wind speed captures the slope of intensification with SST quite well 164 except for very low SSTs (Fig. 3). The explanation for this mismatch at lower SSTs may 165 be that over such cold SST TCs are too weak to compete with random convection and have 166 difficulty in maintaining their mature state. Note that although the theoretical maximum 167 wind speed always exceeds the simulated wind speed, the theoretical pressure perturbation 168 is generally weaker than that in the simulation. Whatever the reason for this difference, the 169 behavior of TCs in the GCM is similar: minimum central surface pressures are comparable 170 to the PI's theory but TCs never reach beyond category 2 in term of maximum wind (Zhao 171 and Held 2010). 172

The sensitivity to ambient rotation rate f is investigated by varying f from 5°N to 20°N 173 with fixed SST at 301 K. Snapshots of surface wind for f at 5°N and 20°N are compared 174 in Fig. 4. As found in HZ08, the number of TCs increases with ambient rotation rate. 175 There are about 23 TCs within the domain for f at 20°N but only about 5 TCs for f 176 at 5°N. The peak frequency of TC central surface pressure shifts towards higher value as 177 f increases, indicating weaker intensity at larger f (Fig. 2b). As f increases from  $5^{\circ}N$ 178 to  $20^{\circ}$ N, the mean central surface pressure (Fig. 3c) increases by  $30 \sim 40$  hPa, and by the 179 same measures the mean maximum surface wind (Fig. 3d) decreases by  $5 \sim 10 \text{ m s}^{-1}$ . The 180 difference in vertical profiles of temperature and specific humidity show a moister and warmer 181 environment for larger ambient rotation rate (Fig. 5). The increasing air temperature and 182

<sup>&</sup>lt;sup>1</sup>utilizing the code provided at ftp://texmex.mit.edu/pub/emanuel/TCMAX/

humidity at the surface corresponds to a reduced surface moisture disequilibrium. The 183 domain-averaged evaporation and precipitation do not change substantially, as expected from 184 the close balance between radiation cooling and latent heat flux, but the average surface wind 185 speed increases due to the increased density of TCs. We interpret the reduction in surface 186 disequilibrium as determined by the need to maintain the same strength of hydrological 187 cycle in the presence of stronger mean surface wind. The increasing of temperature is 188 amplified upward, approximately following a moist adiabat. The pattern of temperature 189 and humidity differences also indicates more aggressive shallow convection in the cases with 190 larger ambient rotation rate. The potential intensity computed from this vertical profile 191 satisfactorily captures the decreasing intensity with increasing f. 192

To study the sensitivity to surface drag coefficient, we bypass the Monin-Obukhov scheme 193 and prescribe surface exchange coefficients homogeneously within the domain. The surface 194 drag coefficient  $C_D$  is varied from  $0.6 \times 10^{-3}$  to  $2.4 \times 10^{-3}$ , while the surface exchange coef-195 ficient for enthalpy  $C_k$  is fixed at  $1.2 \times 10^{-3}$ . Three sets of experiments are conducted with 196 ambient rotation rate set at 10°N, 15°N and 20°N. The distribution of TC central surface 197 pressure is shown in Fig. 6. Interestingly, this curve shifts with  $C_D$  in opposite direction for 198 different rotation rates. For smaller  $C_D$ , there are more intense TCs with f set at 10°N but 199 more weak TCs with f set at  $20^{\circ}$ N. In spite of this difference, the number of TCs within the 200 domain decreases with  $C_D$  in all three cases (as can be seen in Fig. 6). 201

The sensitivity of TC intensity to  $C_D$  is summarized in Fig. 3e and 3f, with ambient 202 rotation rate chosen at 15°N. As  $C_D$  decreases from  $2.4 \times 10^{-3}$  to  $0.6 \times 10^{-3}$ , the mean central 203 surface pressure intensifies by only about 10 hPa and the mean maximum surface wind 204 increases by roughly 50%. The difference in vertical profiles of temperature and specific 205 humidity are shown in Fig 5, indicating a warmer and more humid environment for smaller 206  $C_D$ , similar to the cases with larger f. The potential intensity computed from this vertical 207 profile shows a stronger sensitivity to  $C_D$  compared to the model results, especially for 208 potential pressure intensity. Modifying the ambient rotation rate does not change this result. 209

We note that the recent modification to PI theory by Emanuel and Rotunno (2011) weakens the sensitivity to  $C_D$ .

The mean radius-height structure of the strongest TC within the domain for two SSTs is 212 shown in Fig. 7. The boundary layer inflow and upper-tropospheric outflow characteristic 213 of tropical cyclones is apparent, with an ascent region just inside of the radius of maximum 214 surface wind  $(R_{mw})$ . As shown in Fig. 8,  $R_{mw}$  increases with SST, decreases with  $C_D$  and 215 increases with f for small f (but saturates and probably decreases as f is increased further). 216 The strongest storms that form for a given parameter setting have the smallest  $R_{mw}$ s.  $R_{mw}$ s 217 simulated in this 25-km model are substantially larger than those observed (Stern and Nolan 218 2011). However we feel that the systematic variation seen in this model is of interest for 219 comparison with high resolution simulation. We return to these results in section 4 where we 220 show that quantitatively different results are obtained in smaller domains containing only 221 one storm. 222

HZ08 explored the sensitivity of the natural extent of TCs to SST and f in the turbulent multi-storm equilibrium. With a higher resolution and a larger domain size, we hope to have a better estimate of this sensitivity. As shown in Fig. 9, the average number of TCs, n, doubles or halves as SST is perturbed by 4 K from the reference SST (301 K) and increases nearly linearly with f. We estimate the natural extent of TCs ( $r_0$ ) as

$$r_0 = \sqrt{\frac{A}{n}} \tag{1}$$

<sup>228</sup> in which A is the domain area. Two simple theoretical scalings for the Rossby radius of <sup>229</sup> deformation NH/f, and the other is a natural length scale from PI theory  $V_p/f$ , where  $V_p$ <sup>230</sup> is the potential intensity (Emanuel 1988). However, as shown in Fig. 9, while both of these <sup>231</sup> two scalings project the right trend, neither satisfactorily captures the sensitivity. Modifying <sup>232</sup> some of the choices in the computation of potential intensity, or using actual model intensities <sup>233</sup> for the  $V_p/f$  scalings does not affect this conclusion. We will return to this topic in section <sup>234</sup> 4.

### <sup>235</sup> 4. Varying domain size

The sensitivity of rotating radiative-convective equilibrium to domain size is investigated by varying the domain size from  $(1.25 \times 10^3 \text{ km})^2$  to  $(1.25 \times 10^4 \text{ km})^2$ , with SST at 301 K and f at 10°N. The composites of surface wind for different domain sizes are shown in Fig. 10. When the domain is small, only one tropical cyclone develops. The tropical cyclone has some elasticity in the sense that it can expand to fill the domain until multiple storms finally appear. Particularly, the extent of the single TC in the  $(7.5 \times 10^3 \text{ km})^2$  domain is nearly 20% larger than the natural extent of TCs in the  $(1.25 \times 10^4 \text{ km})^2$  domain.

The behavior of TCs evolves through four regimes, as illustrated by the time series of 243 minimum surface pressure in each simulation (Fig. 11). When the domain is very small, 244 only a tropical depression, with its central surface pressure at about 980 hPa, develops 245 within the domain. As the domain size increases, the depression evolves into an intermittent 246 tropical cyclone with intensity varying widely over a cycle of approximately 50 days. Further 247 increase of the domain size avoids this collapse to weak intensity and leads to a sustained 248 storm. Eventually, if the domain size becomes sufficiently large, multiple storms coexist. 249 (We have conducted preliminary simulations analogous to these with a higher resolution 250 non-hydrostatic model and see similar regime behavior, although the necessary domain size 251 for a sustained storm is somewhat smaller.) 252

To see how the natural extent of TCs in turbulent equilibrium is related to this regime 253 behavior, we conduct two series of experiments with : 1) different f at 5°N, 10°N and 20°N 254 with SST at 301 K, and 2) different SST at 301 K and 305 K with f at 10°N. The time-255 mean central surface pressure is summarized in Fig. 12a and 12b respectively. The intensity 256 first increases with domain size and then levels off approximately at the beginning of the 257 regime of a single sustained storm (marked with an O symbol). This regime shifts towards 258 large domain size as SST increases or f decreases, the same as the natural extent of TCs 259 in turbulent equilibrium. For different parameter settings, the natural extent of TCs in 260 turbulent equilibrium (marked with a | symbol) always falls in this regime. The following 261

simple picture is consistent with these results: If the number of TCs increases from its 262 equilibrium value, the available space for a given TC must decreases into the intermittent 263 storm regime. This TC become unstable, and is more likely to be eliminated by competition 264 with surrounding storms when sufficiently weak. The number of TCs then returns to its 265 equilibrium value. On the other hand, if the number of TCs decreases, the suppression effect 266 of TCs on its surrounding area becomes weak as TCs expand. With fewer but larger TCs 267 randomly moving around, there is more free space for TC genesis. These two effects favor 268 new TC genesis and bring the number of TCs back to its equilibrium. 269

One explanation for this regime of a sustained storm is that the available space for 270 TCs needs to be sufficiently large to provide enough angular momentum to support the 271 inner vortex. In addition to the two scales mentioned in section 3, a third candidate for 272 the natural extent of TCs then comes up by relating the angular momentum at the outer 273 radius  $r_0$  to that at  $R_{mw}$ . For a reduction factor  $\alpha$  between these two, this relationship can 274 be written as  $V_p R_{mw} = \frac{\alpha}{2} f r_0^2$ . Assuming this reduction factor  $\alpha$  is nearly invariant when 275 SST or f varies, we have  $r_0 \sim \sqrt{R_{mw} \frac{V_p}{f}}$ . This is essentially the same relationship between 276  $R_{mw}$  and  $r_0$  as in Emanuel and Rotunno (2011). As shown in Fig. 9, this scaling nicely 277 captures the variation of  $r_0$  with both SST and f except at very low SSTs. As we mentioned 278 previously, the equilibrium achieved at very low SSTs, with a large number of weak TCs, 279 looks very different from those with higher SST. The above argument may only apply to the 280 cases where mature storms dominate. This parameter fit is a good description of the model 281 results. However this relationship between  $r_0$  and  $R_{mw}$  does not in itself provide a theory 282 for  $r_0$  or  $R_{mw}$ . 283

The variation of  $R_{mw}$  with parameters (SST, f and  $C_D$ ) in lattice equilibrium, with domain size at 5000 km, is shown in Fig. 13.  $R_{mw}$  increases with f and decreases with SST and  $C_D$ . This result is consistent with that of Schecter and Dunkerton (2009) in which a three-layer tropical cyclone model is used, but different from what we found in turbulent equilibrium, for SST in particular (Fig. 8). The parameter sensitivity of  $R_{mw}$  in lattice

equilibrium reflects the parameter sensitivity of the ratio between  $R_{mw}$  and TC extent  $r_0$ , 289 since in a fixed size small domain  $r_0$  is fixed, while the parameter sensitivity of  $R_{mw}$  in 290 turbulent equilibrium takes into account both the parameter sensitivity of this ratio and 291 that of  $r_0$ . As we discussed in section 3, TC extent  $r_0$  can vary significantly with parameters 292 in turbulent equilibrium. In this interpretation, the increase of  $r_0$  with SST dominates over 293 the decrease of  $\frac{R_{mw}}{r_0}$  and results in increasing  $R_{mw}$  with SST in turbulent equilibrium (Fig. 294 8a), while the decrease of  $r_0$  with f may offset the increase of  $\frac{R_{mw}}{r_0}$  and leads to saturation 295 or even decrease of  $R_{mw}$  for large f in turbulent equilibrium (Fig. 8b). 296

### <sup>297</sup> 5. Conclusion and Discussion

Idealized models can be used to help fill the gap between theory and comprehensive climate model efforts. In much recent work, the spherical geometry is held fixed while the column physics and/or surface boundary condition are largely simplified. Instead, we choose to retain the column physics and simplify the geometry and surface boundary conditions. By coupling the column physics to the hydrostatic dynamics in a doubly-periodic domain, we can obtain a homogeneous doubly periodic rotating radiative-convective framework.

This study is along the line of Held and Zhao (2008), with the column physics replaced 304 by that of a meso-scale resolution global atmospheric model, HiRAM, which produces quite 305 realistic TC climatology when run with realistic boundary condition. A large domain  $(2 \times 10^4)$ 306  $\mathrm{km}^{2}$  is used to obtain turbulent multi-storm equilibria. The parameter sensitivity is then 307 studied by varying sea surface temperature, ambient rotation rate f and surface drag coef-308 ficient  $C_D$ . The mean TC intensity increases with SST and decreases with f and  $C_D$ . The 309 potential intensity (Bister and Emanuel 2002a) calculated from the thermodynamic environ-310 ment captures the sensitivity to SST and f with reasonable accuracy but overestimates that 311 to  $C_D$ , especially in terms of central surface pressure. Note that PI theory in itself doesn't 312 depend on f so an understanding of the dependence of the intensity on f requires a theory 313

for the sensitivity of the model's equilibrated thermodynamic profiles to f.

The natural extent of TCs,  $r_0$ , in turbulent equilibrium increases with SST and decreases with f. Although we have some difficulty in estimating  $r_0$  by simple theoretical scalings, we find it related to a shift in the regime behavior of lattice equilibrium, from a unsteady to a relatively stable mature storm. Such regime behavior is not unique in our model with meso-scale resolution and hydrostatic dynamics; we have also observed it in preliminary simulations using a high resolution non-hydrostatic model.

In turbulent equilibrium, the radius of maximum surface wind  $(R_{mw})$  increases with 321 SST, increases with f for small f but saturates or probably decreases as f is increased 322 further. However in lattice equilibrium,  $R_{mw}$  decreases with SST and keeps increasing with 323 f. This is because TC extent  $r_0$  is constrained by the domain size in lattice equilibrium 324 while in turbulent equilibrium  $r_0$  varies systematically with environmental parameters. It is 325 the relationship between  $R_{mw}$  and  $r_0$  that is robust to change in domain size. Because of the 326 poor resolution of the  $R_{mw}$  in this model, the results on the parameter sensitivity of  $R_{mw}$ 327 are necessarily tentative. 328

We believe that simulations of rotating radiative-convective equilibrium such as these, with the resolution and column physics of a global atmospheric model being used to study the response of TCs to climate change, can provide a needed middle ground between the global model and non-hydrostatic cloud-resolving simulations in an idealized geometry that will help in establishing the strengths and limitations of the global simulations. This preliminary descriptive study is meant to help encourage further work along these lines.

### 335 Acknowledgments.

The authors thank Ming Zhao and Shian-Jiann Lin for making HiRAM available. We also thank Bruce Wyman for his work in configuring the doubly periodic model. Wenyu Zhou is partly supported by Award NA08OAR4320752 from National Oceanic and Atmospheric Administration (NOAA) Cooperative Institute for Climate Science, and partly by DOE <sup>340</sup> award DE-SC0005189.

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- <sup>397</sup> matology, interannual variability, and response to global warming using a 50-km resolution
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FIG. 1. Snapshots of surface wind (m s<sup>-1</sup>) for different sea surface temperatures. (left) SST = 297 K; (right) SST = 305 K. Ambient rotation rate f is set at 20°N.



FIG. 2. (a) The average number of TCs in the domain with their central surface pressures within the specified 10 hPa interval, as a function of SST. Ambient rotation rate f is set at 20°N. (b) As in (a), but for the sensitivity to ambient rotation rate f. SST is fixed at 301 K. For both sets of experiments, surface exchange coefficients are calculated from Monin-Obukhov scheme.



FIG. 3. (a) The gray line is the median central surface pressure as a function of SST of all collected TCs. The 25th and 75th percentiles coincide with the edges of the gray box. The black line shows the mean of the strongest storms at each instant of time. The red line shows the estimate from potential intensity theory using the domain- and time-mean thermodynamical profile. (b) As in (a), but for maximum surface wind. SST varies from 280 K to 307 K. Ambient rotation rate f is fixed at 20°N. (c), (d) are similar to (a) and (b), but for the sensitivity to ambient rotation rate f. SST is fixed at 301 K. For both sets of experiments, surface exchange coefficients are calculated from Monin-Obukhov scheme. (e), (f) show the sensitivity to surface drag coefficient  $C_D$ . Surface drag coefficient  $C_D$  varies from  $0.6 \times 10^{-3}$  to  $2.4 \times 10^{-3}$  with surface exchange coefficient for enthalpy  $C_k$  fixed at  $1.2 \times 10^{-3}$ . Ambient rotation rate f is fixed at 15°N and SST is 301 K.



FIG. 4. Snapshots of surface wind (m s<sup>-1</sup>) for different ambient rotation rates. (left) f at 5°N; (right) f at 20°N. SST is fixed at 301 K.



FIG. 5. (a, b) Differences in vertical profiles of temperature and specific humidity between different ambient rotation rates (vs 5°N case) with surface exchange coefficients calculated from Monin-Obukhov scheme. and (c, d) surface drag coefficients (vs  $C_D = 2.4 \times 10^{-3}$  case) with ambient rotation rate at 15°N. SST is fixed at 301 K.



FIG. 6. (a) The average number of TCs in the domain with their central surface pressures within the specified 10 hPa interval as a function of surface drag coefficient  $C_D$  (0.6, 1.2,  $2.4 \times 10^{-3}$ ) with ambient rotation rate f set at 10°N. (b) (c) As in (a), but with ambient rotation rate f set at 15°N and 20°N. Surface exchange coefficient for enthalpy  $C_k$  is fixed at  $1.2 \times 10^{-3}$  and SST is 301 K.



FIG. 7. Time-mean and azimuthal-mean tangential (colors), radial (white) and vertical (black) velocities of the strongest tropical cyclone for sea surface temperature a) 297 K and b) 305 K. Tangential winds are contoured every 5 m s<sup>-1</sup>. Positive (cyclone) values are solid, negative (anticyclone) values are dashed, and the zero contour is a red dashed line. Radial winds are contoured at [2,5,10,15,20] m s<sup>-1</sup>. Positive (outflow) values are dashed, negative (inflow) values are solid, and the zero contour is omitted. Vertical velocity is contoured (negative only) at -1, -2, and -3 Pa  $s^{-1}$ . Ambient rotation rate f is fixed at 20°N.



FIG. 8. (a) The gray line is the median  $R_{mw}$  of all collected TC as a function of SST. The 25th and 75th percentiles coincide with the edges of the gray box. The black line shows the mean  $R_{mw}$  of the strongest storms at each instant of time. SST varies from 280 K to 307 K. Ambient rotation rate f is fixed at 20°N. (b) as in (a), but for the sensitivity to ambient rotation rate f. SST is fixed at 301 K. For both sets of experiments, surface exchange coefficients are calculated from Monin-Obukhov scheme. (c) are those for the sensitivity to surface drag coefficient  $C_D$ . Surface drag coefficient  $C_D$  varies from  $0.6 \times 10^{-3}$  to  $2.4 \times 10^{-3}$  with surface exchange coefficient for enthalpy  $C_k$  fixed at  $1.2 \times 10^{-3}$ . Ambient rotation rate f is fixed at 15°N and SST is 301 K.



FIG. 9. (a) The average number of TCs n (black-dash line with diamond) and the natural extent of TCs  $r_0$  ( $\sqrt{\frac{A}{n}}$ , gray dash-dot line with circle) as a function of SST. Ambient rotation rate f is set at 20°N. Three possible scalings are shown. Red line:  $\sqrt{R_{mw}\frac{V_p}{f}}$ , discussed in section 4; blue line:  $\frac{V_p}{f}$ ; green line:  $\frac{NH}{f}$ . (b) As in (a) but for the variation of ambient rotation rate f. The underlying SST is 301 K. The estimates from these three scalings are normalized to the natural extent of TCs with SST at 301 K and f at 20°N.



FIG. 10. Composite of surface wind for simulations with domain sizes at 1250 km, 2500 km, 5000 km, 7500 km and snapshot of surface wind for domain size at 12500 km, all with SST at 301 K and f at 10°N.



FIG. 11. Time series of minimum surface pressure for different regimes. SST is 301 K and ambient rotation rate f is fixed at 10°N.



FIG. 12. Time mean TC central surface pressure as a function of domain sizes at (a) different ambient rotation rates and (b) sea surface temperatures. The four regimes mentioned above are marked with different symbols,  $\nabla$  (tropical depression), + (intermittent storm),  $\circ$  (sustained storm) and  $\triangle$  (multiple storms). The average scale of TCs in corresponding large domain simulation is marked with a | symbol.



FIG. 13. Sensitivity of  $R_{mw}$  to f (black line), SST (red line) and  $C_D$  (blue line) in a fix size small domain (5000 km on a side). In f and  $C_D$  perturbation experiments, SST is fixed at 301 K. For both SST and  $C_D$  perturbation, two series of experiments with ambient rotation rates set at 10°N (dash-dot line) and 20°N (dash line) are conducted.